

2 ASTROBIOLOGY  
0 GRADUATE  
1 CONFERENCE  
7



CHARLOTTESVILLE, VA



1  
00:00:00,790 --> 00:00:07,210

[Music]

2  
00:00:11,780 --> 00:00:09,620

for those of you who don't know me I'm

3  
00:00:14,990 --> 00:00:11,790

Jay crawl I also go by Alec that's how

4  
00:00:16,760 --> 00:00:15,000

you'll find me on social media I use

5  
00:00:20,000 --> 00:00:16,770

masculine gender pronouns he has in him

6  
00:00:21,650 --> 00:00:20,010

and I study I'm a graduate student at

7  
00:00:24,320 --> 00:00:21,660

the University of Colorado where I study

8  
00:00:26,859 --> 00:00:24,330

one how gender and sexuality affect the

9  
00:00:29,990 --> 00:00:26,869

academic outcomes for Gen chem students

10  
00:00:32,150 --> 00:00:30,000

more topically I also study sulfur

11  
00:00:34,700 --> 00:00:32,160

reactions and atmospheres in planetary

12  
00:00:36,020 --> 00:00:34,710

atmospheres I'm not going to talk to you

13  
00:00:38,270 --> 00:00:36,030

about either of those things this

14

00:00:46,270 --> 00:00:38,280

morning I'm going to talk to you about

15

00:00:53,860 --> 00:00:52,069

rocks cool so more specifically I'm

16

00:00:55,850 --> 00:00:53,870

going to talk to you a little bit about

17

00:00:57,590 --> 00:00:55,860

some of the rocks that you're going to

18

00:00:59,380 --> 00:00:57,600

hear about in this morning's session and

19

00:01:01,819 --> 00:00:59,390

why we care about them as

20

00:01:05,000 --> 00:01:01,829

astrobiologists to the geologists in the

21

00:01:10,070 --> 00:01:05,010

room please forgive me and correct me

22

00:01:11,270 --> 00:01:10,080

during the session after so the first

23

00:01:14,090 --> 00:01:11,280

rock that we're going to talk about is

24

00:01:16,670 --> 00:01:14,100

Venus it's often referred to as Earth's

25

00:01:18,649 --> 00:01:16,680

twin more accurately it might be

26

00:01:22,550 --> 00:01:18,659

described as our evil twin on a bad hair

27

00:01:25,760 --> 00:01:22,560

day but it's very similar to earth in

28

00:01:27,530 --> 00:01:25,770

size and mass and Composition but a

29

00:01:29,780 --> 00:01:27,540

couple of things make it very notably

30

00:01:31,749 --> 00:01:29,790

different especially at sulfuric acid

31

00:01:35,899 --> 00:01:31,759

clouds so from about 50 to 70 kilometers

32

00:01:39,440 --> 00:01:35,909

there's a global sulfuric acid cloud

33

00:01:42,200 --> 00:01:39,450

layer it also has an incredibly slow day

34

00:01:45,230 --> 00:01:42,210

so it takes about 243 Earth days for

35

00:01:46,730 --> 00:01:45,240

Venus to turn once on its axis and it

36

00:01:49,520 --> 00:01:46,740

also rotates in the opposite direction

37

00:01:51,289 --> 00:01:49,530

of all the other planets this is likely

38

00:01:56,450 --> 00:01:51,299

because of some massive cataclysmic

39

00:01:58,880 --> 00:01:56,460

collision or earlier in its history it

40

00:02:03,800 --> 00:01:58,890

also is incredibly hot at its surface

41

00:02:06,200 --> 00:02:03,810

about 500 degrees Celsius and if we take

42

00:02:08,210 --> 00:02:06,210

a look at its atmosphere that's a little

43

00:02:10,760 --> 00:02:08,220

bit easier to explain the surface

44

00:02:12,470 --> 00:02:10,770

pressure is 93 bar it's almost 100 times

45

00:02:12,830 --> 00:02:12,480

more pressure at its surface than on

46

00:02:15,559 --> 00:02:12,840

earth

47

00:02:18,050 --> 00:02:15,569

and that atmosphere is almost entirely

48

00:02:19,550 --> 00:02:18,060

carbon dioxide and so it has this

49

00:02:22,880 --> 00:02:19,560

incredibly potent greenhouse gas

50

00:02:26,270 --> 00:02:22,890

atmosphere there's some nitrogen as well

51  
00:02:30,229 --> 00:02:26,280  
as some other trace gases most notably

52  
00:02:32,150 --> 00:02:30,239  
so2 but because of that global cloud

53  
00:02:35,449 --> 00:02:32,160  
cover it has this incredibly high albedo

54  
00:02:37,759 --> 00:02:35,459  
which means that almost 90% of the light

55  
00:02:39,920 --> 00:02:37,769  
that is in coming from the Sun actually

56  
00:02:41,509 --> 00:02:39,930  
gets reflected away from Venus but the

57  
00:02:44,420 --> 00:02:41,519  
atmosphere is so thick that even that

58  
00:02:46,970 --> 00:02:44,430  
10% that gets in and absorbed by the

59  
00:02:50,420 --> 00:02:46,980  
planet leads to incredible heating of

60  
00:02:52,460 --> 00:02:50,430  
the surface there's also very little

61  
00:02:56,300 --> 00:02:52,470  
hydrogen and Venus's atmosphere and

62  
00:02:58,610 --> 00:02:56,310  
there's an incredibly high d2h ratio and

63  
00:03:00,259 --> 00:02:58,620

so what that tells us is that actually

64

00:03:03,229 --> 00:03:00,269

Venus has probably lost a lot of the

65

00:03:05,210 --> 00:03:03,239

water from it which is interesting

66

00:03:07,039 --> 00:03:05,220

because early on its history we think

67

00:03:10,210 --> 00:03:07,049

Venus may have been a lot more like

68

00:03:12,830 --> 00:03:10,220

Earth and had oceans on its surface

69

00:03:18,020 --> 00:03:12,840

today its surface looks incredibly

70

00:03:21,650 --> 00:03:18,030

different though there is because of its

71

00:03:23,330 --> 00:03:21,660

similar size and density we assume that

72

00:03:26,539 --> 00:03:23,340

it's pretty similar to earth and that it

73

00:03:28,280 --> 00:03:26,549

has a core a mantle and a crust however

74

00:03:30,259 --> 00:03:28,290

we don't have a lot of actual direct

75

00:03:33,500 --> 00:03:30,269

evidence about the internal structure of

76

00:03:34,910 --> 00:03:33,510

the planet we don't see any evidence of

77

00:03:36,920 --> 00:03:34,920

plate tectonics which is pretty

78

00:03:41,690 --> 00:03:36,930

interesting and there's no magnetic

79

00:03:44,420 --> 00:03:41,700

field however most of the impact craters

80

00:03:47,660 --> 00:03:44,430

are almost in pristine condition the

81

00:03:49,039 --> 00:03:47,670

surface is very young only about 300 to

82

00:03:51,080 --> 00:03:49,049

600 million years old

83

00:03:54,080 --> 00:03:51,090

and there's been a lot of evidence of

84

00:03:56,420 --> 00:03:54,090

past volcanic activity and so what we

85

00:03:59,660 --> 00:03:56,430

think happens on Venus is that the crust

86

00:04:02,390 --> 00:03:59,670

is doesn't we don't have any of these

87

00:04:04,099 --> 00:04:02,400

plate tectonics right now but the mantle

88

00:04:06,080 --> 00:04:04,109

slowly heats up and at some point

89

00:04:08,090 --> 00:04:06,090

there's enough to actually enough energy

90

00:04:10,280 --> 00:04:08,100

to actually disrupt the crust and then

91

00:04:14,740 --> 00:04:10,290

we see massive turnover of the surface

92

00:04:17,090 --> 00:04:14,750

and resurfacing of the entire planet

93

00:04:19,009 --> 00:04:17,100

with so if we go a little bit farther

94

00:04:21,670 --> 00:04:19,019

out in our solar system we've got Titan

95

00:04:24,469 --> 00:04:21,680

it's the largest moon orbiting Saturn

96

00:04:25,980 --> 00:04:24,479

but perhaps most interesting is that

97

00:04:30,120 --> 00:04:25,990

it's the only moon with a

98

00:04:34,189 --> 00:04:30,130

atmosphere and the surface is composed

99

00:04:37,320 --> 00:04:34,199

mostly of water ice and rocky material

100

00:04:39,300 --> 00:04:37,330

but this dense atmosphere is possibly an

101  
00:04:40,350 --> 00:04:39,310  
analogue to early Earth's atmosphere

102  
00:04:44,040 --> 00:04:40,360  
which is what makes it really

103  
00:04:46,620 --> 00:04:44,050  
interesting to us as astrobiologists and

104  
00:04:48,990 --> 00:04:46,630  
so its atmosphere actually has a very

105  
00:04:50,460 --> 00:04:49,000  
similar structure to Earth it has this

106  
00:04:53,129 --> 00:04:50,470  
troposphere where we see all of its

107  
00:04:54,960 --> 00:04:53,139  
weather and then an upper stratosphere

108  
00:04:58,260 --> 00:04:54,970  
where there's a stolen haze which is

109  
00:05:00,629 --> 00:04:58,270  
mostly carbon containing organics with

110  
00:05:03,240 --> 00:05:00,639  
some nitrogen in them and then a miso

111  
00:05:05,939 --> 00:05:03,250  
sphere and thermosphere its atmosphere

112  
00:05:09,180 --> 00:05:05,949  
is almost 90 almost entirely nitrogen

113  
00:05:10,350 --> 00:05:09,190

it's about 98% nitrogen we do see some

114

00:05:13,080 --> 00:05:10,360

methane

115

00:05:15,060 --> 00:05:13,090

it's 1.2 percent overall but as you go

116

00:05:18,089 --> 00:05:15,070

closer to the surface actually increases

117

00:05:20,909 --> 00:05:18,099

to about 5% of the atmosphere and that's

118

00:05:22,620 --> 00:05:20,919

because it's incredibly cold and so that

119

00:05:24,540 --> 00:05:22,630

methane actually begins to freeze out

120

00:05:28,550 --> 00:05:24,550

and you get precipitation of it on the

121

00:05:31,680 --> 00:05:28,560

planet in this lower troposphere and

122

00:05:33,210 --> 00:05:31,690

there's a little bit of hydrogen and

123

00:05:35,010 --> 00:05:33,220

then there's a lot of trace organics

124

00:05:40,550 --> 00:05:35,020

formed because of the photo chemistry

125

00:05:42,930 --> 00:05:40,560

that goes on with methane before Cassini

126  
00:05:46,350 --> 00:05:42,940  
cassini-huygens we actually did not know

127  
00:05:48,060 --> 00:05:46,360  
a whole lot about Titan surface

128  
00:05:50,040 --> 00:05:48,070  
thanks to that mission we actually do

129  
00:05:52,740 --> 00:05:50,050  
now know that there are hydrocarbon

130  
00:05:54,300 --> 00:05:52,750  
lakes and rivers and that the surface is

131  
00:05:56,070 --> 00:05:54,310  
actually pretty geologically young

132  
00:05:57,870 --> 00:05:56,080  
somewhere between a hundred and a

133  
00:06:01,379 --> 00:05:57,880  
billion 100 million and a billion years

134  
00:06:04,890 --> 00:06:01,389  
old but we can actually see these canals

135  
00:06:07,170 --> 00:06:04,900  
and rivers on the surface as well as a

136  
00:06:11,070 --> 00:06:07,180  
number of lakes at the northern pole of

137  
00:06:13,290 --> 00:06:11,080  
the moon as well and then also

138  
00:06:16,020 --> 00:06:13,300

particularly interesting it has a rocky

139

00:06:17,760 --> 00:06:16,030

silicate core about 3,400 kilometers in

140

00:06:20,430 --> 00:06:17,770

diameter so it's a little bit bigger

141

00:06:23,790 --> 00:06:20,440

than Earth's moon but smaller than the

142

00:06:26,790 --> 00:06:23,800

earth but there's a lot of evidence that

143

00:06:30,089 --> 00:06:26,800

there is a subsurface ocean possibly a

144

00:06:32,070 --> 00:06:30,099

global ocean and what we think would

145

00:06:34,980 --> 00:06:32,080

allow for this is actually an ammonia

146

00:06:36,899 --> 00:06:34,990

water mixture so once you have enough

147

00:06:40,019 --> 00:06:36,909

ammonia in that water you can get this

148

00:06:45,060 --> 00:06:40,029

eutectic mixture that doesn't freeze out

149

00:06:46,379 --> 00:06:45,070

until much colder and so that also gives

150

00:06:48,680 --> 00:06:46,389

us the potential for a lot of

151  
00:06:53,279 --> 00:06:48,690  
interesting chemistry and importance for

152  
00:06:54,390 --> 00:06:53,289  
astrobiology the next rock that we're

153  
00:06:57,209 --> 00:06:54,400  
probably going to hear about is

154  
00:06:59,579 --> 00:06:57,219  
meteoroids and the IAU definition is

155  
00:07:01,909 --> 00:06:59,589  
incredibly specific it's a solid object

156  
00:07:04,620 --> 00:07:01,919  
moving in the interplanetary space of

157  
00:07:06,570 --> 00:07:04,630  
considerable size considerably smaller

158  
00:07:11,579 --> 00:07:06,580  
than an asteroid and considerably larger

159  
00:07:13,920 --> 00:07:11,589  
than an atom and so this actually gives

160  
00:07:16,730 --> 00:07:13,930  
us a pretty wide berth but most of these

161  
00:07:20,370 --> 00:07:16,740  
things are about 4.5 billion years old

162  
00:07:22,079 --> 00:07:20,380  
and perhaps most interestingly about 15

163  
00:07:24,420 --> 00:07:22,089

thousand metric tons of these are

164

00:07:28,140 --> 00:07:24,430

delivered to earth every year and so

165

00:07:30,510 --> 00:07:28,150

these are a source of incoming material

166

00:07:33,689 --> 00:07:30,520

that can seed the planet with a lot of

167

00:07:37,079 --> 00:07:33,699

interesting material over you know the

168

00:07:38,879 --> 00:07:37,089

course of billions of years additionally

169

00:07:42,329 --> 00:07:38,889

most of them come from the asteroid belt

170

00:07:44,579 --> 00:07:42,339

about 99.8% of meteorites that we've

171

00:07:46,409 --> 00:07:44,589

found have come from there and so they

172

00:07:50,300 --> 00:07:46,419

can tell us a lot about the early solar

173

00:07:52,649 --> 00:07:50,310

system composition and some is well our

174

00:07:54,089 --> 00:07:52,659

cometary debris and so that's normally

175

00:07:57,480 --> 00:07:54,099

what we would refer to is like a meteor

176

00:07:59,640 --> 00:07:57,490

shower and then a very small number of

177

00:08:01,469 --> 00:07:59,650

them are ejected material from

178

00:08:04,469 --> 00:08:01,479

collisions with much larger body such as

179

00:08:06,829 --> 00:08:04,479

Mars in the moon and so these are some

180

00:08:11,189 --> 00:08:06,839

of the first hints that we got at what

181

00:08:15,749 --> 00:08:11,199

composition what the composition of Mars

182

00:08:17,159 --> 00:08:15,759

and the moon were after that we'll hear

183

00:08:19,589 --> 00:08:17,169

about some stromatolite some

184

00:08:20,879 --> 00:08:19,599

stromatolites are a little bit different

185

00:08:23,040 --> 00:08:20,889

than all the other rocks we'll be

186

00:08:24,930 --> 00:08:23,050

talking about one there here on earth

187

00:08:26,969 --> 00:08:24,940

which is exciting

188

00:08:30,240 --> 00:08:26,979

but they're built by living microbial

189

00:08:31,740 --> 00:08:30,250

colonies of cyanobacteria and on the

190

00:08:35,430 --> 00:08:31,750

modern earth they're actually found in

191

00:08:37,860 --> 00:08:35,440

only a very few select places such as

192

00:08:39,209 --> 00:08:37,870

Shark Bay and in the Caribbean and this

193

00:08:40,980 --> 00:08:39,219

is because they have to form in very

194

00:08:45,270 --> 00:08:40,990

hyper saline conditions and I'll talk a

195

00:08:46,829 --> 00:08:45,280

little more about that here but in the

196

00:08:47,429 --> 00:08:46,839

Precambrian area they were fairly

197

00:08:49,860 --> 00:08:47,439

ubiquitous

198

00:08:51,300 --> 00:08:49,870

all over the planet and they may date

199

00:08:54,059 --> 00:08:51,310

back all the way back to the Ark

200

00:08:55,590 --> 00:08:54,069

so there's some of the earliest forms of

201  
00:08:57,480 --> 00:08:55,600  
life that we see on the planet that we

202  
00:09:01,230 --> 00:08:57,490  
actually have like really nice records

203  
00:09:03,749 --> 00:09:01,240  
of and so the way these things form

204  
00:09:07,019 --> 00:09:03,759  
cyanobacteria are these photosynthetic

205  
00:09:09,090 --> 00:09:07,029  
species so they need to be able to get

206  
00:09:14,069 --> 00:09:09,100  
sunlight to grow but you get these mats

207  
00:09:16,290 --> 00:09:14,079  
of cyanobacteria and then they form on

208  
00:09:20,280 --> 00:09:16,300  
the shallow waters the deepest been

209  
00:09:21,840 --> 00:09:20,290  
found about 14 meters or 45 feet but as

210  
00:09:24,269 --> 00:09:21,850  
they do that they build up this mucus

211  
00:09:27,239 --> 00:09:24,279  
layer that can trap sediment or they can

212  
00:09:29,939 --> 00:09:27,249  
precipitate calcite as well but as that

213  
00:09:32,040 --> 00:09:29,949

happens then they grow up through that

214

00:09:33,900 --> 00:09:32,050

layer and they start a new mat and then

215

00:09:35,489 --> 00:09:33,910

they grow out and then they collect more

216

00:09:36,960 --> 00:09:35,499

sediment and then they get another layer

217

00:09:38,160 --> 00:09:36,970

and they grow up through it and then

218

00:09:41,100 --> 00:09:38,170

they start another layer and they do

219

00:09:43,619 --> 00:09:41,110

this over and over again and so they

220

00:09:45,269 --> 00:09:43,629

trap a lot of the interesting material

221

00:09:47,340 --> 00:09:45,279

so they tell us a lot about what the

222

00:09:50,160 --> 00:09:47,350

conditions were when they were growing

223

00:09:54,210 --> 00:09:50,170

what was available what was settling out

224

00:09:55,650 --> 00:09:54,220

in the oceans but it's important that

225

00:09:58,470 --> 00:09:55,660

they need to be free of any sort of

226  
00:10:00,869 --> 00:09:58,480  
grazing species so if anything eats them

227  
00:10:02,999 --> 00:10:00,879  
then they can't build these rocks that's

228  
00:10:05,040 --> 00:10:03,009  
very problematic in today's oceans

229  
00:10:07,139 --> 00:10:05,050  
however pre-cambrian there weren't any

230  
00:10:09,389 --> 00:10:07,149  
fish or anything eating them so we have

231  
00:10:10,590 --> 00:10:09,399  
a lot of them around and you get them in

232  
00:10:12,960 --> 00:10:10,600  
all sorts of really interesting

233  
00:10:18,449 --> 00:10:12,970  
structures from domes to columns to

234  
00:10:20,340 --> 00:10:18,459  
splitting columns and then the last Rock

235  
00:10:22,410 --> 00:10:20,350  
we'll call it a rock iron-sulfur

236  
00:10:25,769 --> 00:10:22,420  
clusters think of them as very tiny

237  
00:10:27,720 --> 00:10:25,779  
rocks that are important in biology but

238  
00:10:30,569 --> 00:10:27,730

they're made out of unsurprisingly iron

239

00:10:32,220 --> 00:10:30,579

and sulfur and they come in a really

240

00:10:34,019 --> 00:10:32,230

wide variety of different structures

241

00:10:36,150 --> 00:10:34,029

depending on how much iron and sulfur

242

00:10:39,210 --> 00:10:36,160

you have in them anywhere from very

243

00:10:44,030 --> 00:10:39,220

simple one in two atoms of iron and

244

00:10:46,470 --> 00:10:44,040

sulfur up to a much larger structures

245

00:10:48,449 --> 00:10:46,480

but interestingly they're found in a

246

00:10:49,889 --> 00:10:48,459

number of proteins in biology and

247

00:10:51,929 --> 00:10:49,899

there's at least three biosynthetic

248

00:10:54,210 --> 00:10:51,939

systems that have been identified for

249

00:10:58,199 --> 00:10:54,220

making these things in our biological

250

00:11:00,030 --> 00:10:58,209

systems and what's really interesting is

251  
00:11:03,030 --> 00:11:00,040  
they play a really important role in

252  
00:11:04,710 --> 00:11:03,040  
redox oxidation reactions in

253  
00:11:06,869 --> 00:11:04,720  
mitochondria

254  
00:11:08,819 --> 00:11:06,879  
both complex 1 & 2 an oxidative

255  
00:11:11,939 --> 00:11:08,829  
phosphorylation have multiple iron

256  
00:11:14,160 --> 00:11:11,949  
sulfur clusters in them they can also

257  
00:11:17,220 --> 00:11:14,170  
act as catalytic centers for a lot of

258  
00:11:19,400 --> 00:11:17,230  
biological activity and because of this

259  
00:11:22,499 --> 00:11:19,410  
they've been suggested as playing an

260  
00:11:27,059 --> 00:11:22,509  
important role in the origin of life on

261  
00:11:30,720 --> 00:11:27,069  
Earth and so with that I will thank you

262  
00:11:32,790 --> 00:11:30,730  
guys for your time and hand things over

263  
00:11:34,230 --> 00:11:32,800

to the experts that are going to be

264

00:11:38,690 --> 00:11:34,240

giving talks dare I call them the rock